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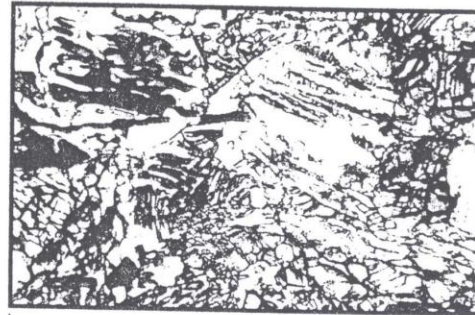
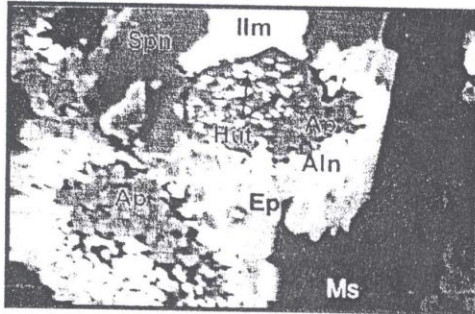
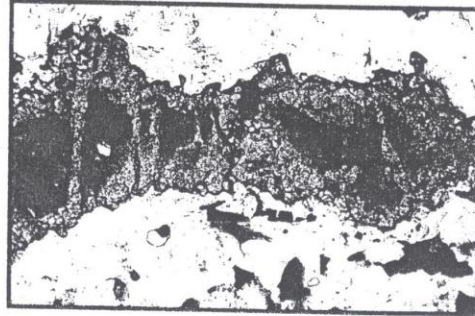
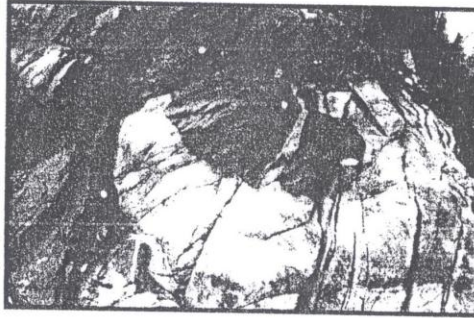
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पैन-अफ्रीकन घटनाक्रम भारत एवं अन्टार्कटिका Pan-African Event India and Antarctica



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A type Malani Magmatism : Signatures of the Pan-African Event In the Northwest Indian Shield Assembly of the Late Proterozoic Malani Supercontinent

Naresh Kochhar

Centre of Advanced Study in Geology, Punjab University, Chandigarh
email : nareshkochhar2003@yahoo.com

Abstract : The plume related Malani magmatism is intra-plate, anorogenic, A-type and is indicative of extensional tectonic environment in the Trans-Aravalli block (TAB, west of the Aravalli- Delhi fold belt) of the NW Indian shield. The Malani supercontinent comprises TAB of the NW Indian shield, Seychelles-Madagascar, Nubian Arabian shield and Yangtze craton of South China. In this paper similarities between these micro continents in terms of bimodal anorogenic magmatism, ring structures, Strutian glaciation and subsequent dessication and discussed. The paleomagnetic data also support the existence of the Malani supercontinent.

Keywords : Malani supercontinent, A-type granite, Rajasthan, Mantle Plume.

INTRODUCTION

The Indian subplate is composed of three geologically different blocks or terranes. The South India Block (SIB), the Bundelkhand Block (BB) and the Trans-Aravalli Block (TAB) which were juxtaposed and sutured during different periods of Earth's history (Radhakrishna, 1989). The Bundelkhand and the South Indian blocks came together during Mid Proterozoic along Central Indian Tectonic Zone (CITZ), the latter has remained active from Early Proterozoic to Recent. The CITZ forms a belt in an ENE-WSW direction and is marked by the Son-Narmada lineament and other associated prominent lineaments. The TAB and BB which are geologically unrelated to each other are separated by NE-SW trending 700km long Proterozoic Aravalli-Delhi mobile belt. (Fig. 1).

Pan African Event

Pan-African Thermo-Tectonic Event : Kennedy (1984) defined the Pan-African as a thermo tectonic event at 500Ma affecting large areas encircling the West African, Congo and Kalahari cratons. It is now in widespread use in other continents and refers to events centered around but not restricted to period from 750-550Ma (Black and Liegeois normal. In South America the term 'Brasilian' is approximately equivalent to the Pan-African (Rogers *et al.*, 1995). The available

geochronological data indicate that Pan-African tectono-thermal event at 500-550Ma is more pronounced in the Lesser and Higher Himalaya and South Indian block, south of Palaghat, Cauvery shear zone and in Wannai Complex and High Land Complex of Sri Lanka. The Himalayan magmatism is of S-type, whereas in Sri Lanka and Tamilnadu the magmatism has A-type affinity. In the Trans-Aravalli block and South Indian block of the Indian shield event at 650-750 Ma is well documented. This magmatism is of A-type (Kochhar, 2001).

In this paper significance of A-type magmatism of Rajasthan and Haryana is discussed in terms of anorogenic and 'within plate' setting, and the assembly of Late Proterozoic Malani supercontinent.

Trans-Aravalli Block (Rajasthan and Haryana)

The TAB is unique in the geological evolution of the Indian shield as it marks a major period of anorogenic (A-type), 'Within Plate', high heat producing (HHP) magmatism represented by the Malani igneous suite of rocks (MIS). The Neoproterozoic Malani igneous suite (55,000 km²; 732 Ma) comprising peralkaline (Siwana), metaluminous to mildly peralkaline (Jalor), and peraluminous (Tusham and Jhunjhunu) granites with cogenetic carapace of acid volcanics (welded

tuff, trachyte, explosion breccia and perlite etc.) and characterized by volcano-plutonic ring structures and radial dykes. The suite is bimodal in nature with minor amounts of basalt, gabbro and dolerite dykes. The Siwana ring structure (30 km in EW 25 km in NS direction) is the most spectacular feature of the Thar desert. The representatives of Malani suite also occur at Kirana Hills, and at Nagar Parkar, Sindh, Pakistan (Qasem Jan et.al, 1997) and is shown

in Fig. 1. The Malani magmatism is controlled by NE-SW trending lineaments (zones of extension and high heat flow) of fundamental nature (mantle) and owes its origin to mantle plume (for a review, Kochhar, 2000a). Recently Bonin (2007) has suggested that A-type ring complex granites were produced in planetary environment differing from that prevailing on Earth. Rare felsic material found in the meteoric and lunar record yield dominantly A-type granites.

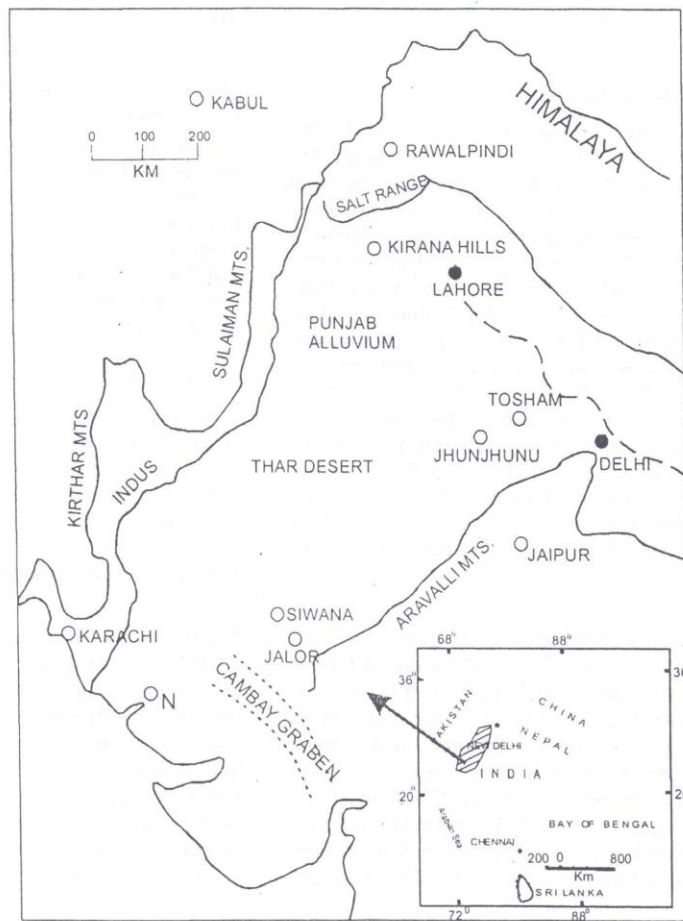


Fig. 1 : Location map of the Malani Igneous Suite S : Siwana, J : Jalor, JH : Jhunjhunu, SD : Sarnu-Dandali, N : Nagarparkar

Evidences for Precambrian mantle plume activity in the TAB

Low velocity anomaly

Kennette and Widiwantoro (1999) based on P wave arrival time delineated a low velocity anomaly to the north of Cambay Gulf. This feature is 120 km across and is in marked contrast to high seismic velocities which characterize the lithosphere beneath the peninsular India. This anomaly extends from shallow depth to contact with more extensive low velocity zone below 200 km beneath the Indian lithosphere. The anomaly coincides with the Siwana ring structure, and the low velocity anomaly may represent the conduit of a fossil plume, the Malani plume (Kochhar, 2001).

Gravity and heat flow data

The Marwar terrane of the TAB is characterized by high heat flow and basement high indicating underplating and uplift. The Tusham area is associated with the highest heat flow of 96 mWm^{-2} in the TAB. The heat flow observed in the around Khetri and Jhunjhunu is 75 mWm^{-2} , with an average value of 60 mWm^{-2} for the Delhi fold belt. High heat flow value of 93 mWm^{-2} has been reported from Cambay basin.

Chemical and thermal anisotropy

The chemical and thermal anisotropy in the TAB is manifested in the anorogenic magmatism represented by high heat producing A-type, 'Within Plate'. Malani magmatism indicative of extensional tectonic environment, crustal thickness and high heat flow. The bimodal nature of Malani magmatism is exemplified by the occurrence of basalt, gabbro and dolerite dykes of continental alkaline affinity. The trace element pattern is of high abundance of Hf-Zr, Nb, Ga, Zn, Y, REEs (except Eu) of granites and the associated volcanic rocks emphasize the role of halogens in fluxing these elements from the

mantle (Kochhar and Dhar, 1993; Kochhar, *et al.*, 1995; Kochhar, 2000b; Bhushan, 2000).

Malani Supercontinent

There is a relationship between mantle plume related anorogenic magmatism and assembly of a supercontinent. Plumes initiate continental break up by doming and rifting and rupturing (Morgan, 1972). Preceding Pangea (Late Paleozoic Supercontinent) was the supercontinent of Rodinia generally regarded as persisting to the latest Proterozoic but with wide uncertainties concerning its configuration and times of assembly and dispersal (Rogers *et al.*, 1993). According to Dalziel (1992) breaking up of Rodinia occurred during two phases one at 750 Ma and the younger rifting phase at 550 Ma. The breakup at 750 Ma resulted in the separation of East Gondwana from the present west coast of Laurentia.

Spot paleomagnetic readings at 750 Ma also suggest that neither India nor Australia can be joined in either a traditional East Gondwana or Rodinia fit. The data also does not support the idea of coherent East Gondwana (Meert, 2001) Anekt-Mozambique ocean closed to form Arabian – Nubian shield – the join between East and West Gondwana during 870-640 Ma (Unrug, 1992, Rogers *et al.*, 1993).

The period ca. 732 Ma mark a major Pan-African tectono-magmatic event of widespread magmatism, of alkali granites and comagmatic acid volcanics (anorogenic, A-type) in the Trans – Aravalli block of the Indian shield, Central Iran (Forster, 1987), Nubian – Arabian shield (Kroner *et al.*, 1989) and Madagascar and South China (Yoshida *et al.*, 1999) Kochhar, 2007a, b, Li *et al.*, 1999, 2004); Somalia (Kroner *et al.*, 1990) and Seychelles Hoshino, 1986 Kochhar, 2004).

In view of the widespread ca. 750Ma alkaline magmatism, which is so widespread, and well

Table 1: Similarities between TAB, Nubian Arabian Shield, Seychelles, and South China in terms of bimodal, anorogenic magmatism, protolith, paleopoles and Strutian glaciation and subsequent desiccation

	TAB NW INDIAN SHIELD	SEYCHELLES	NUBIAN ARABIAN SHIELD	S. CHINA YANGTZE BLOCK
Alkali Magmatism granites	Siwana Jalor Tusham ca 745 Ma	Mahe, Ste Anna Praslin ca. 700-755 Ma	Jabal-Al Hasasin Jabal-AS Sawal Dahul Ca. 700-760 Ma	Chengjian ca. 780-745 Ma
Protolith	Archean BGC	Archean BGC	Paleo Proterozoic or older crustal Precursor	-
Paleopole	55-70 ^U	30 ^U N	-	70 ^U N
Strutian glaciation	Pokhran, Salt Range Boulder Bed	-	Haqf Super Group	Linntua and Nantua deposits
Evaporite/ Carbonate Sequence	Hanseran, Nagaur Bikaner Basin	-	Hormuz Series	Late Sinian Late Precambrian carbonate and phosphates.

developed in TAB, central Iran, Arabian – Nubian shield, Madagascar and S.China, Somalia, Seychelles, Kochhar (1996, 2000a,b, 2001, 2007) proposed that all these microcontinents were characterized by a common crustal stress pattern, rifting and thermal regime, Shrutian glaciation and dessication and similar paleolatitudinal positions which could be attributed to the existence of a supercontinent – the Malani Supercontinent, (Table 1) (Fig. 2).

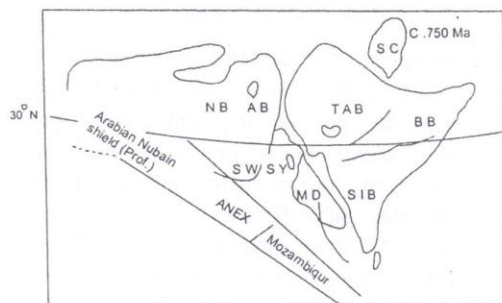


Fig. 2 : Assembly of the Malani supercontinent, TAB : Trans-Aravalli Block, BB : Bundelkhand Block, SIB : South Indian Block, NB-AB : Nubian-Arabian Shield, SM : Somalia, SY : Seychelles, MD : Madagascar, SC : South China

Seychelles

The Mahe (750Ma) and Ste. Anne (764 Ma) granites of Seychelles have been correlated with 732 Ma old hypersolvus (Siwana) and subsolvus (Jalor granites) of MIS. These granites have A-type signatures and owe their origin to the same protolith of Archean crust of (amphibolite) BGC of Rajasthan as indicated by Sm/Nd; Rb/Sr isotopic data. The Seychelles and Malani granites have similar low oxygen isotope ratios $\delta^{18}\text{O}$ SMOW (+3 TO +4‰ (Seychelles granites) and Siwana and Jalor granite) (-0.10 to +1.8‰, -4.60 to +1.2‰) (Ashwal *et al.*, 2002, Kochhar, 2004).

Similarities between granites of Mis (Jalor and Siwana) and the Seychelles (Mahe, STE. Anne and Praslin) Granites : (Kochhar, 2004)

1. The coeval (730 Ma) (Rb/Sr) hypersolvus (Siwana) and subsolvus (Jalor) granites of MIS (Kochhar and Dhar, 1993) have been correlated with Mahe (750 Ma) and Ste. Anne (764 Ma) granites of Seychelles (Ashwal *et al.*, 2002).

2. Role of Archean basement granites complex crust (2.5-2.3 Ga) in the evolution of granites of both the regions. Pb and Nd isotopic compositions of the Siwana granites show that the magma is mantle derived, and for Jalor complex, the combined Sr and Nd data indicate primary mantle derivatives with a variable degree of contamination crust of Archean age (Dhar *et al.*, 1996). According to Ashwal *et al.*, (2002) the Archean crust of BGC Rajasthan, 3.2 Ga of amphibolitic composition with initial Sr and Nd isotopic composition at 750 Ma could be suitable source of Mahe granitic rocks of Seychelles.
3. Tectonic discrimination diagrams: Ga/Al vs. Zr plot of Mahe and Praslin granites indicate A-type affinity, and the data overlap with the fields of Jalor and Tusham granites of MIS. Similarly majority of Seychelles samples plot in the WPG field of Ta vs Yb and Nb vs Y diagrams of Pearce *et al.*, (1984). It is surprising that Ashwal *et al.*, (2002) did not plot these tectonic discriminating diagrams.
4. Oxygen isotopic data : The Seychelles granites are characterized by low ^{18}O (+3 to +4) ‰ as compared to +5 to +11‰ values of most terrestrial igneous rocks. This corresponds well with the low ^{18}O values of +1.0 to 1.8‰ of Siwana granites and -4.0 to 1.21 values of Jalor granites (Kochhar, 2000a).

Torsvic *et al.*, (2001a) obtained a paleomagnetic pole with latitude 74.5°N and longitude 71.2°E at 750 Ma for the MIS, whereas, Klootwijk (1975) has obtained 81°S latitude and 22.4°E longitude for MIS at 729 Ma. Paleomagnetic data place the MIS at 41°N at local paleolatitude when combined with normal polarity site of Klootwijk (1975) and the Seychelles at 30°N (Torsvic *et al.*, 2001). At 750 Ma B.P. Seychelles and MIS were only 600 km apart. Since the 'Within Plate'.

Anorogenic nature and extensional tectonic environment of MIS is established, its correlation with Mahe, Praslin and Ste. Anne granites is suggestive of similar setting for the Seychelles granites. Stephens *et al.*, (1997) also favour intraplate, extensional tectonic environment for Seychelles granites which could be related to the plume tectonics. The MIS does not represent Andean type arc on the western margin of Rodinia.

Madagascar

Within central block of Madagascar, Carion granites which have alkaline geochemical signatures and are post tectonic were emplaced during extensional collapse of Pan-African Orogen U/Pb ages of granites of western central Madagascar near Ambistra indicate emplacement age of 804-775 Ma, some of these granites have gabbroic rocks as sleeves (Handke *et al.*, 1999). In this respect they are very much similar to Jalor granites.

The Arabian Shield

The Arabian Nubian shield (ANS) is an accreted terrane composed of crustal domains of island arc, oceanic and continental affinities overlain by post accretionary sediments and volcanics, and invaded by voluminous intrusions. The Saudi Arabia part of the Nubian Arabian shield has been divided into five terranes separated by four suture zones. The three terranes i.e. Asir, Hijaz, Midyan west of the Nabitah suture zone are of ensimatic character, whereas, eastern two i.e. Afif and Ar-Rayan are of continental to marginal continental character. The three ensimatic terranes are collectively referred to Arabian-Nubian arc terranes (Stoesser, 1986).

Most of the shield was created in two distinct stages. These are (1) Island arc stage (between 660-900 Ma) in which subduction related processes operated for the generation of magma, and (2) a post accretionary stage (between 550-700 Ma) where magma was generated by fusion of continental protolith (Stoesser, 1986).

According to the Jackson (1986) the post accretionary stage is marked by the presence of granitic intrusions. According to Stern *et al.* (2004) the evolution of the Arabian-Nubian shield occurred between 870 Ma and the end of Precambrian (ca. 542 Ma) and the crustal growth encompassed a time of dramatic climate change (see below).

In southern Yemen, the amalgamation of the two terranes i.e. Abas (west) and Al-Mahfid (east), separated by Al Bayda island arc sequence, is marked by the 760 Ma magmatic activity (Whitehouse *et al.*, 2004). According to Johnson and Kattan (2004) the Asir terrane is more closely linked structurally and tectonically with East African orogen, Madagascar, Ethiopia, Somalia, Yemen and Eritrea than with the northern Shield, and that the southern and northern Arabian shields may be decoupled.

The remnant of the Archean crust as Proterozoic basement has been reported on the basis of lead isotope studies in the Afif terrane. A zircon from trondjemite along the Al Amar fault has yielded 2067 ± 74 Ma age there by indicating Late Proterozoic crust in the shield. Granodiorite gneiss from Jabal Khida in the easternmost part of the shield has been dated at 1629 ± 20 Ma (Stoesser, 1986).

According to Stoesser *et al.*, (2004) there is no indication of crust formation between 1670-900 Ma as indicated by U/Pb and single zircon data. Between 760-740 Ma ago, the Khida terrane separated from the margin of Paleoproterozoic and Archean continental block-the Arabian craton, presently concealed beneath the Phanerozoic cover in the east but exposed in Yemen.

Magmatism

The Arabian shield contains one of the largest fields of alkali granites in the world. Forty nine major and more than a dozen minor alkali granite plutons occur in the Midyan and Hijaz terranes and in the Nabitah orogenic belt. The alkali granites do not occur SW and easternmost part of the shield. The alkali granites were the last

intrusive phase and were emplaced during a span of 180 Ma between 680-510 Ma ago. The alkali granites are spatially associated with metaluminous hypersolvus to subsolvus granites, the latter forming the core of complexes. The alkali granites are also associated with a swarm of alkali rhyolite dykes in the northeastern part of the shield (Stoesser and Elliot 1985). A total of thirty three ring complexes and funnel shaped intrusions have been recognized related to the extensional tectonic regime in the Arabian shield. (Fig. 3) The magmatism (670-570 Ma) is mainly peralkaline to metaluminous with A-type affinity. The granites and the related rock types were emplaced by cauldron subsidence which operated in the final stage of the cratonisation of the Arabian shield (Roobal and White, 1986).

The peraluminous and peralkaline granites with A-type geochemical characteristics are encountered at Jabal-Al Hasasin, Jabal As Sawal, Jabal Dahul (Stoesser *et al.*, 2004). According to Vail and Kuron (1978) high level ring complexes, ring dykes and cone sheets have been reported from the Red Sea Hills and Bayuda desert in NE Sudan. These complexes comprise granites, quartz syenite, nepheline syenite gabbro etc, and have been dated at 700 Ma. The Abas terrane is dominated by 760 Ma granitoids. (Whitehouse *et al.*, 2004), with Sm-Nd model ages (2.3 Ga) indicating reworking of a Paleoproterozoic or older crustal precursor. These features are very much similar to the Siwana and Jalor granites of MIS (Kochhar, 2006).

Interestingly Forster (1987) compared the Infra-Cambrian and Cambrian rhyoliths and the associated granites of Central Iran with the Malani rhyoliths and the Pan-African granites of Arabian shield. The similarity has been attributed to the Gondwana connection based on paleomagnetic data. Central Iran held a position south of equator during Infra Cambrian and later drifted from the southern margin of the Gondwanaland to its present position as a microplate. According to Kroner *et al.*, (1990), the period between 850-750 Ma represents the most significant crust formation event in all parts

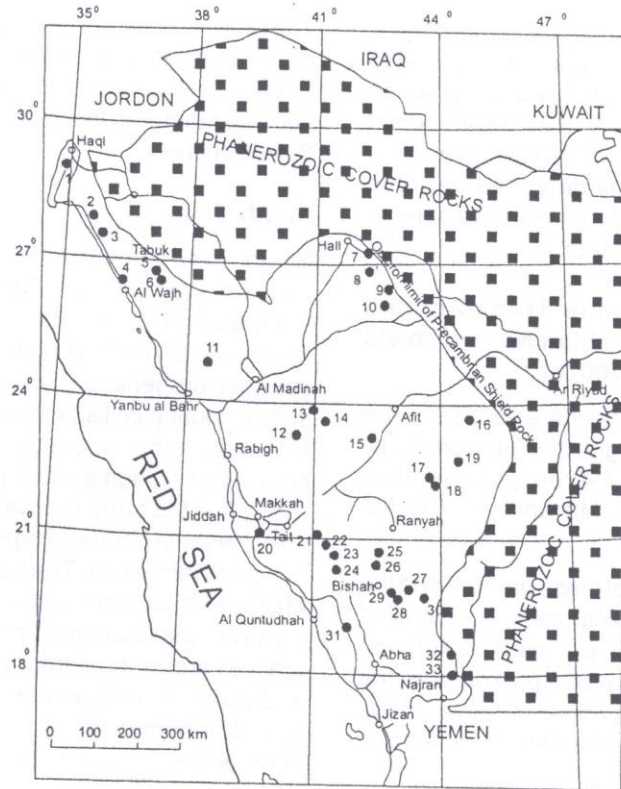


Fig. 3 : Map showing the location of Ring complex (RC), Single ring structure (SR), Beejar plutons (BJP) and Calderas (C) in the Arabian shield

Sr. No.	Location	Type	Geochemical Affinity
1	Sawda Complex	RC	AA
2	Dabbagh Complex	RC	AA
3	Sulay Siyah Complex	RC	P
4	Liban pluton	BJP	P
5	Habd Complex	BJP	P
6	Warid Complex	RC	AA
7	Jabal Salma Complex	C	AA
8	Jabal AL-Makhrugah	SR	P
9	Jabal Raha	C	AA
10	Jabal Silrilah	SR	AA
11	Martabah Complex	RC	P
12	Jabal Ramram	C	P
13	Jabal Hadash Shara	RC	P
14	Jabal Hadbad Dayahm	RC	AA
15	Jabal as Safwah	RC	AA
16	Jabal Arawan	RC	AA
17	Jabal Kursh	RC	AA

Sr. No.	Location	Type	Geochemical Affinity
18	Jabal Dahul Complex	RC	AA
19	Uyaijah ring dyke	RC	AA
20	Unnamed Complex	RC	P
21	Jabal Al-Qunnah	RC	P
22	Unnamed ring dyke	SR	AA
23	Khutamah ring dyke	SR	P
24	Aquq ring dyke	SR	P
25	Jabal Munirah	RC	P
26	Jabal Al Najjah	RC	P
27	Jabal Bani-Bisqan	RC	AA
28	Jabal Al-Hasser	RC	AA
29	Jabal Refdah	RC	P
30	Unnamed Complex	RC	P
31	Lakathah Complex	RC	P
32	Wadi Idimah Complex	RC	AA
33	Jabal Ashirah Complex	RC	AA

AA : Alkaline affinity, syenite, alkali granite with riebeckite, aegirine etc. A-type. P Peraluminous / metaluminous affinity

of ANS including Somalia in the form of felsic intraplate magmatism. Rogers (1990) has also suggested that both Indian shield and Arabian Nubian shields are similar in their development including production of alkali granites, subsidence of thick partly deformed basin on recently formed crust and ultimate development of platformal cover sediments.

Similarities between the Trans-Aravalli Block (TAB) and the Arabian Nubian shield (ANS) (Kochhar, 2006, 2007b)

1. Both the terranes are characterized by the occurrence of alkali granites (peraluminous-peralkaline with A-type geochemical signatures) at 732 Ma and ca. 700 Ma respectively.
2. The emplacement of these granites and the associated acid volcanics was controlled by ring structures and cauldron subsidence. Ring complexes are the continental manifestations of plume activity and are indicative of extensional tectonic environment.
3. Both the terranes preserve the remnants of Archean crust i.e. BGC of amphibolitic facies (3.2 Ga in the TAB (Kochhar, 2004) and trondjemite of 2.6 Ga in the Afif terrane of ANS. This older crustal protolith has acted as precursor to the younger alkali granites.
4. In both the terranes the alkali granites mark the cratonisation of the shield.
5. Both the terranes are characterized by similar peralkaline and peraluminous granite hosted mineralisation of U, Th, Nb, Ta REEs, Sn, W, Li and F etc. In this respect they are akin to the younger granites of Nigeria. (Kochhar, 1998, 2000a and b, Vallinayagam, 1999 and Somani 2006).
6. Both the terranes are characterized by the occurrence of evaporites-carbonate sequence and show evidence of Strutian glaciation.

The aforesaid similarities can only be explained if the Arabian Nubian shield was attached to the Trans-Aravalli block of the Indian shield around 600-700 Ma B.P. in the configuration of the Malani Supercontinent.

South China

South China comprises the YC to the northeast and the Precambrian Cathaysia block to the southeast. The YC has as Late Archaean-Proterozoic nucleus surrounded mostly by younger orogenic belts. Both the blocks were sutured during collision of Grenville age (Fig. 4). Neoproterozoic anorogenic magmatism and coeval mafic magmatism are widespread around the YC of South China. The magmatism coincides with NS-trending Kangdian rift and NE-trending Nanhua rift. The first one at ca. 830-795 Ma (pre-rift) and the second one at 780-745 Ma (syn-rift), the Chengjian magmatism. The later granites are younger than 1.0 Ga Sibao orogeny and intrude the rift sequence. Both the pre- and syn-rift magmatism have been attributed to superplume because of its intraplate setting, peralkaline to alkaline affinity and association with coeval swarms of mafic dykes. Emplacement of the dyke swarm initiated the break-up of Rodinia. The superplume may represent two normal plumes to account for these two episodes of magmatism. According to Wang and Li (2001), it places the ca. 780-750 Ma granites are unconformably overlain by Upper Sinian sequence (750 Ma), which is interpreted to have formed a rift cover. Li *et al.* (2004) have correlated the syn-rift (780-745 Ma) magmatism with the Malani magmatism besides those of Laurentia, South Africa and Australia.

Paleomagmatism and Strutian glaciation

The APWP palaeomagnetic poles at ca. 800 Ma place South China at 55-70°N at par with high paleolatitudes of India during Malani rhyolite period (Ca. 750Ma) (Li *et al.*, 2004, Torsvic *et al.*, 2001). The migration of India to higher paleolatitudes could be the cause of Precambrian glaciation as exemplified by the Pokhran boulder in Rajasthan which is of glaciogenic origin

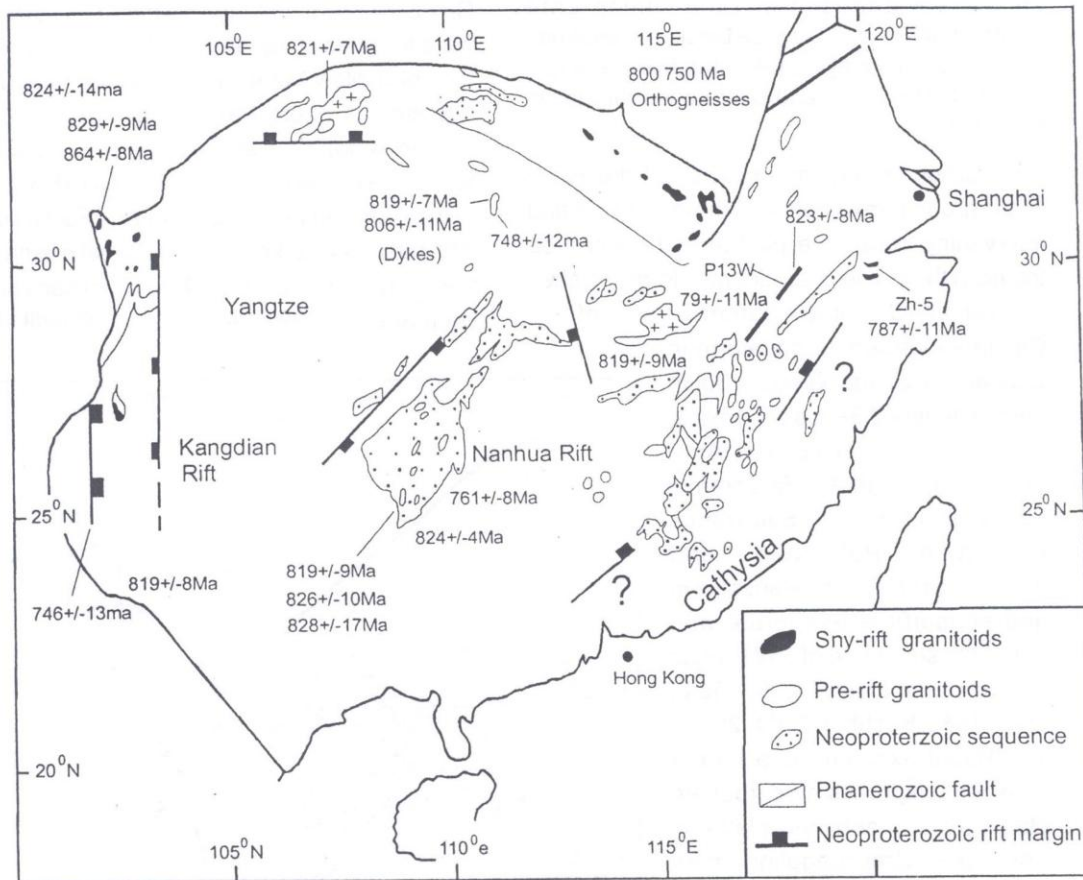


Fig. 4 : Neoproterozoic tectonic framework of South China New ages are highlighted bold

(Chauhan *et al.*, 2001, Pareek, 1984 and Sinha-Roy *et al.*, 1998). Incidentally, the Pokhran diamictite correlates with other Precambrian diamictites in the lesser Himalaya such as Blaini boulder bed, Manjir and Bhimdasa in Jammu and Kashmir, Tankaki in Hazara, and Buxa Group (Shergaon pebbled bed) in Arunachal Pradesh (Valdiya, 1995, Virdi, 1995, Tiwari, 2003). Interestingly, Bose *et al.* (1997) have reported acid pyroclastics such as welded tuff, ash flow tuff, lapilli tuff, volcanic clasts from the Late Proterozoic formations of Lesser Himalaya, viz. Shimla Group, Blaini Formation, Infrakrol Formation and the Krol Group. The source of these pyroclastics could be attributed to the nearby Tusham acid volcanics of Malani age. It

is possible that the precursor to the Lesser Himalyan terrane was contiguous part of the TAB during Late proterozoic. The boulders of Malani granites also occur in the Salt Range. In the YC, Liantua and Nautua (748 Ma) deposits are of glaciogenic origin. Around 700Ma Bp India started drifting towards the tropical latitudes and around 600 Ma, India was near the equator. According to Pooranchandra Rao *et al.*, (1997) there has been large polar movement between the Malani rhyolite and other lower Cambrian period poles indicating rapid migration of the Indian subcontinent from the northern to southern hemisphere latitudes. During this period, the northern margins of India were subjected to desiccation. The carbonate, mainly dolomite and

phosphate deposits of Late Sinian-Late Precambrian age in Yangtze block, corresponds with Hanseran evaporates in the Marwar basin of TAB, SW Punjab, and Bilara phosphorites in Rajasthan.

The carbonate evaporate facies of the Haqf Supergroup form a belt of evaporate basin and intervening carbonate platform which can be traced in these micro-continents: Hormuz Series of Arabian Gulf and southern Iran, Infra – Cambrian carbonate platform in Central Iran, Ghabar Group of South Yemen, India Ad Series of Somalia, Wadi Fatima Series and Abba Formation, South Arabia Saramuj Formation of Jordan, Salt Range Formation, Pakistan, and evaporites of Nagaur-Bikaner basin and subsurface evaporate and anhydrite sequence of SW Punjab and Haryana, India (Fig. 5) (Gorin *et al.*, 1982, Kochhar, 2004, 2006). The Huaqf axis formed a barrier between evaporites an carbonate platform and or anhydrite platform. The Oman-Ural megalineament controlled the distribution of evaporite-carbonate sequence.

Recently, Singh *et al.*, (2006) have described Neoproterozoic evaporite sequence from Lesser Himalaya comprising gypsum and anhydrite, gypsum + carbonate, and halite from the Kashmir sub-basin (Ramban, Assar belt), and Chamba subbasin (Bathri) respectively, and halite from the Guma-Drang area of Mandi, Himachal Pradesh. According to Illyin 1990, the continental margins of Vendian-Early Precambrian super-continental comprising South

China, Kazakhstan, Mongolia and proto-Tethys were the sites of equatorial sedimentation in the form of shallow water carbonate and phosphorite deposits. A short-lived positive excursion in carbon isotope values noticed at phosphorite level in the Birmania (Upper Carbonate) Formation in the TAB and in the basal part of Tal Group in Krol basin is also known from Dahal member of the Yuhucun Formation of Meishucunian zone-II (Kumar *et al.*, 1997). According to Pandit *et al.*,

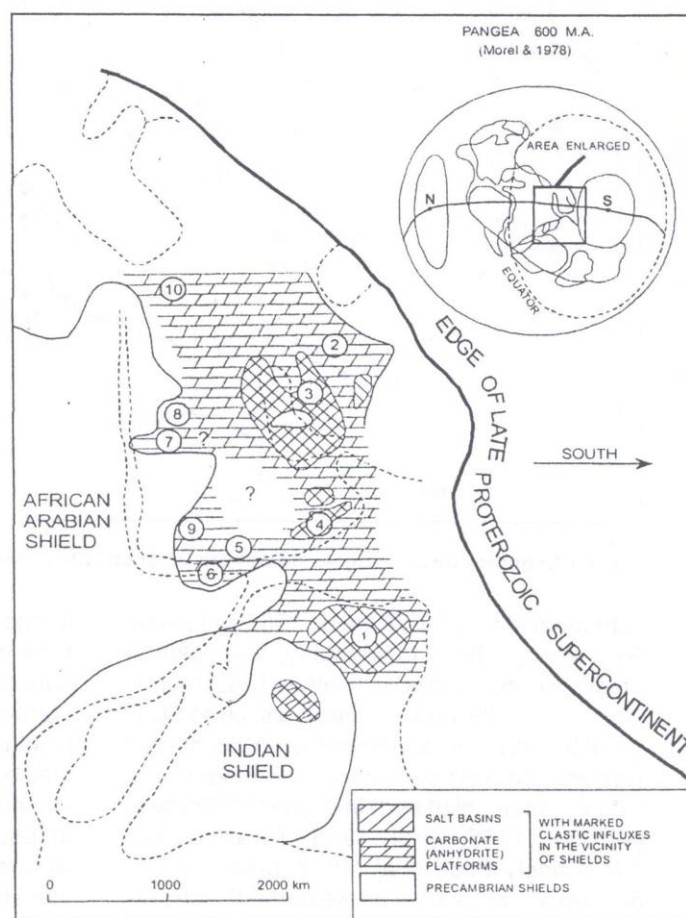


Fig. 5 : Paleogeography of the Late Precambrian-Early Cambrian sediments 1. Salt Range, Pakistan. 2. Zagros Crush Zone, Central Iran, 3. Harmuz Series, Arabia, 4. Haqf group, Oman, 5. Ghabar group, S. Yemen, 6. India Ad Series, Somalia, 7. Wadi Fatima Series, S. Arabia, 8. Wadi Fatima Series, Nougara, S. Arabia, 9. Abba Formation, N. Yemen, 10. Saramuj Formation, Jordan, SW Punjab, Haryana, Rajasthan NW India.

(2001), extremely low C_{13} values in the Bilara carbonate indicate glacial-related cold climatic conditions, while positive shift in the carbon isotope values in the upper formations implies a warmer climatic condition. The Bilara carbonate group carbon isotope profile have close correspondence with global carbon isotopic evolution curve in Haqf, Oman, Siberia platform, Mongolia and Morocco and elsewhere. Li *et al.* (2004) have proposed a 90° spin of Rodinia that brought the entire supercontinent into equatorial latitudes. The spin was the result of initiation of a subequatorial superplume by ca. 750 Ma. The spin corresponds with reversal of geomagnetic field during the deposition of each group of Vindhyan Supergroup, i.e. Kaimur, Rewa and Bhandar. During the Kaimur period (1400-750 Ma) there have been at least three reversals of geomagnetic field (Pooran-chandra *et al.*, 2003). The combined 800-700 Ma apparent polar wander path implies pole to equator rapid velocities of 20 cm/yr for India and South China, which is unlike the other continents such as Australia and Congo. (Kochhar 2001) suggested that since the Indian subplate is a mosaic of three major tectonostratigraphic blocks with different magmatic and metamorphic histories, the position of India in the assembly of a Late Proterozoic supercontinent should not be viewed as a single entity. For example, the southern tip of the SIB (south of Palaghat Cauvery shear zone), Sri Lanka and Madagascar along with the Eastern Antarctica formed a supercontinent (Idnurm and Radhakrishna, 1999). In other supercontinent assembly Meghalaya lies SW of Yilgam craton of Australia, Bungar hills and Wind Mill Islands in an Antarctic continuation of Frazer mobile belt. Due to similarities of bimodal, anorogenic magmatism in South China, India and Australia, Li *et al.*, (2003) have suggested that the superplume was responsible for the break-up of supercontinent during Neoproterozoic. The

Malani plume was responsible for the separation of TAB from East Gondwana. The Malani plume may be the present day position of subequatorial plume at 750 Ma. B.P. proposed by Li *et al.* (2004) (Fig. 6). In view of the occurrence of widespread bimodal, anorogenic plume-related magmatism, Strutian glaciation and similar palaeolatitude position, the YC of South China fits well in the reconstruction of the Malani supercontinent. The position of South China between Australia and Laurentia is debatable because of an alternate equatorial Rodinia supercontinent, excluding high-latitude cratons in India and South China (Li *et al.*, 2003). Whether the Precambrian glaciation in India and South China was due to the flight of these terranes to high latitudes during the Malani rhyolite period or the glaciation was of low latitude due to increased CO_2 drawdown and global albedo affected by waning of plume volcanism is debatable. The evidences presented here



Fig. 6 : Paleogeographic reconstruction of Rodinia 'in tact' at 750 Ma

supported the idea of polar Strutian glaciation in the TAB and Lesser Himalayan terrane. (Kochhar, 2007a).

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